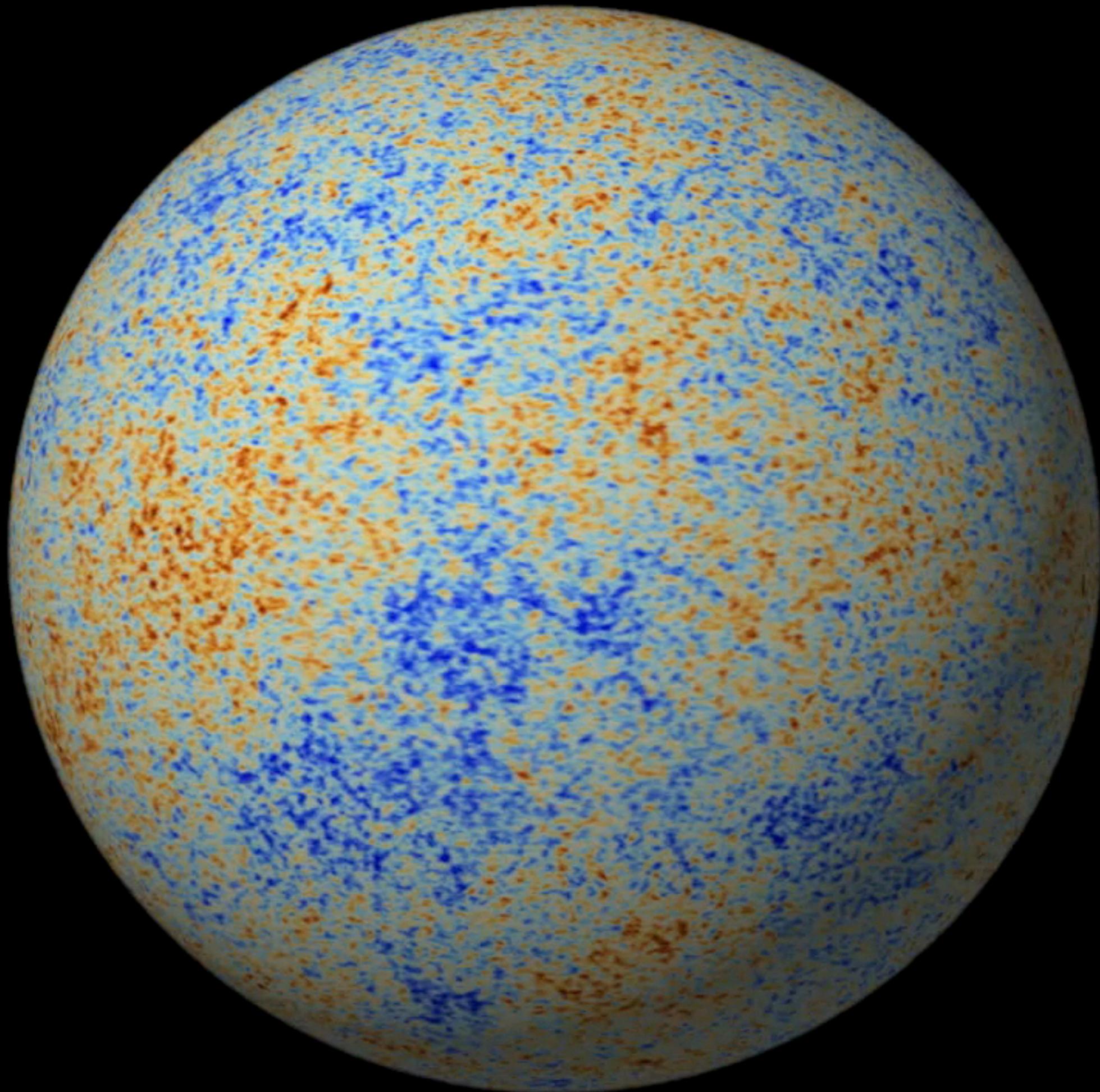


CMB-ish surveys

Douglas Scott

October 2018



Planck papers

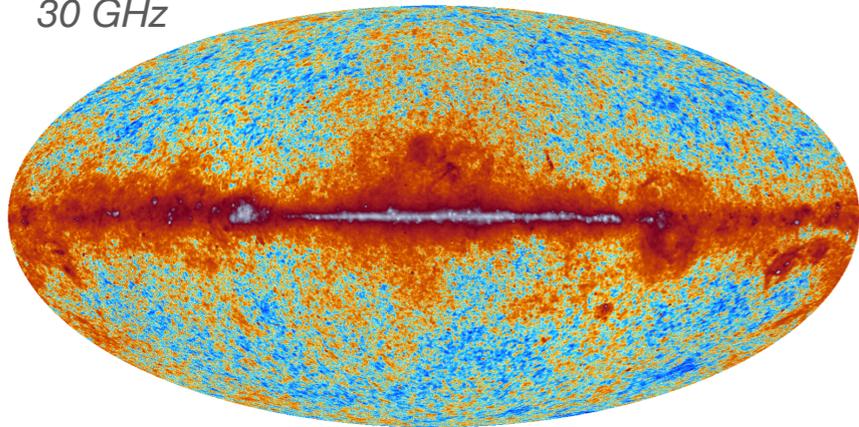
- “Technical papers”
 - 13 papers
- “Planck pre-launch status”
 - 13 papers
- “Planck early results”
 - 26 papers
- “Planck intermediate papers”
 - 54 papers
- “Planck 2013 results”
 - 32 papers
- “Planck 2015 results”
 - 28 papers
- “BICEP & Planck”
- “Planck 2018 results”
 - 13 papers (last 3 very soon!)

Instrumentation
Data processing pipeline
Simulations
Cosmic rays
Zodiacal emission
Component separation
Interstellar medium
Molecular clouds
Galactic magnetic fields
All-sky optical depth
Galactic cold clumps
SNRs & PNe
Anomalous microwave emission
Polarized dust radiation
All-sky CO map
Nearby galaxies
Radio sources & blazars
High-z extragalactic sources
Cosmic infrared background
CMB power spectra
Cosmological parameters
Gravitational lensing
Dipole & boosting effects
Integrated Sachs-Wolfe effect
Reionization
All-sky thermal SZ map
SZ cluster cosmology
Cluster physics
Peculiar velocities
Constraints on inflation
Topological defects
Non-Gaussianity
Isotropy & statistics
Geometry & topology
Dark energy & modified gravity
Primordial magnetic fields
Variation of fundamental constants
Parity & birefringence
Planet fluxes

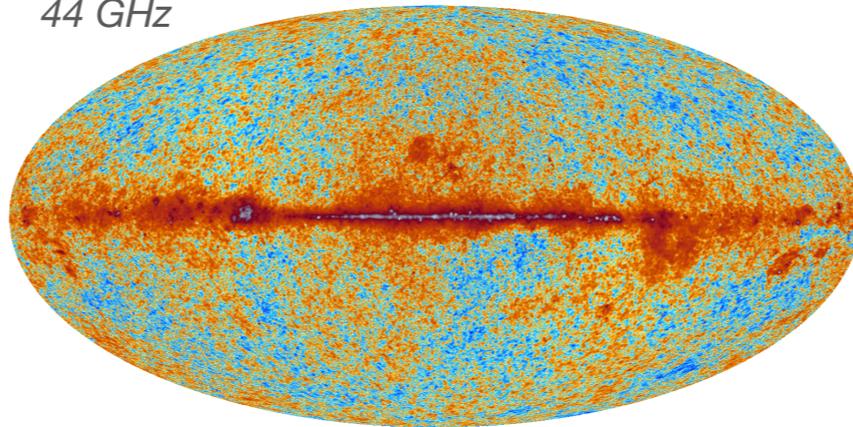
...

All-sky maps of temperature/intensity at 9 separate frequencies

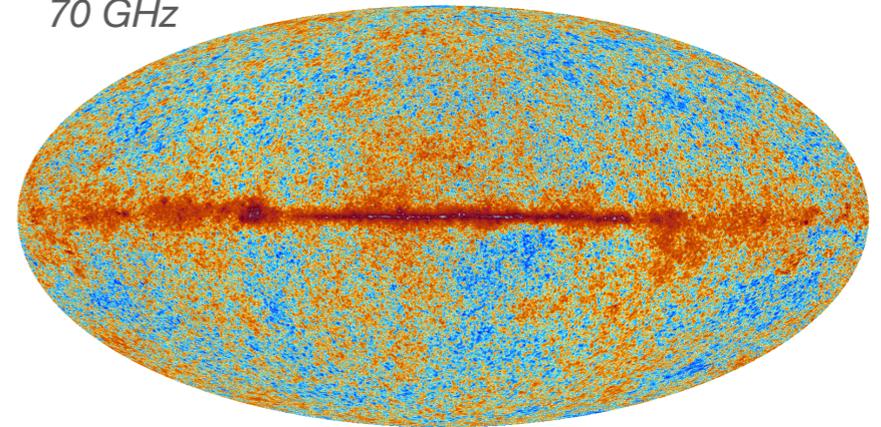
30 GHz



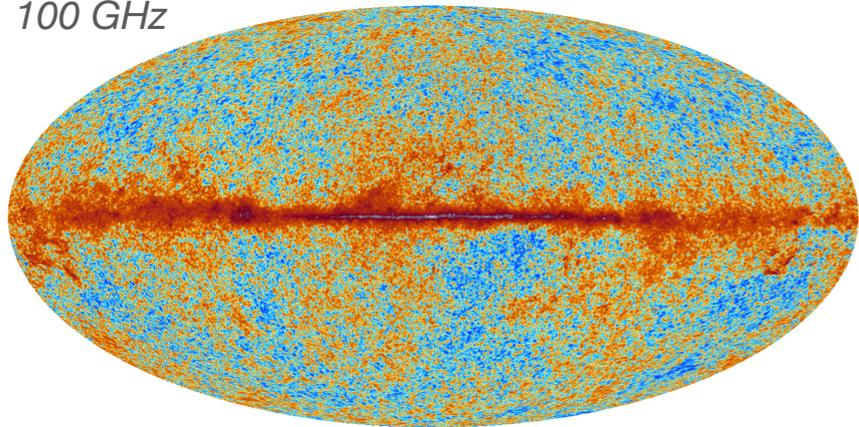
44 GHz



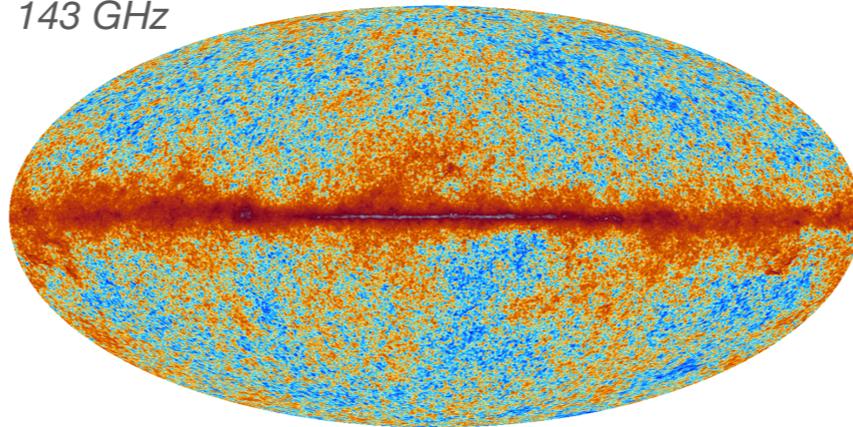
70 GHz



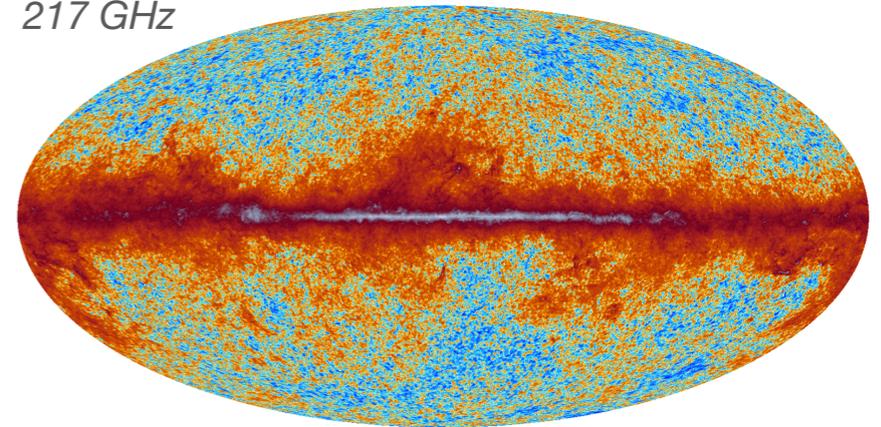
100 GHz



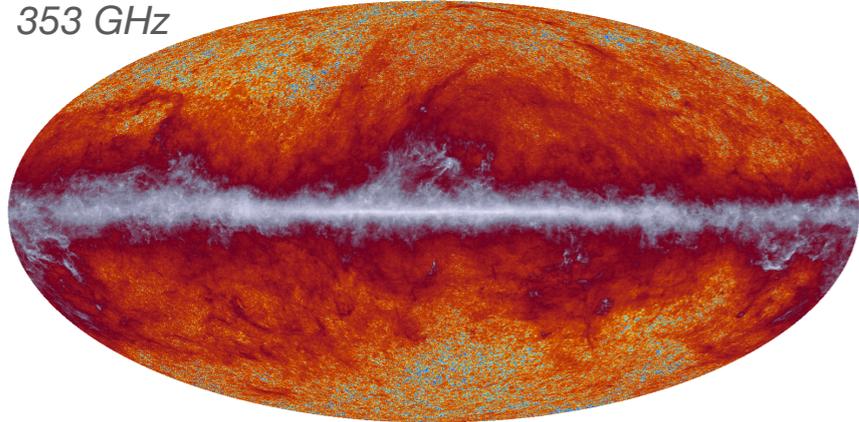
143 GHz



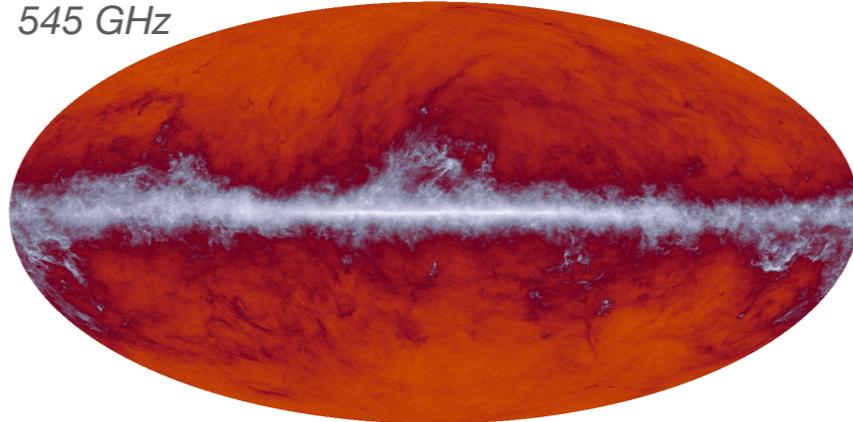
217 GHz



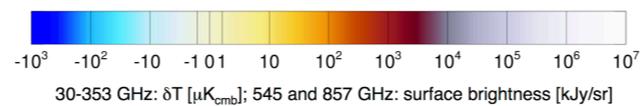
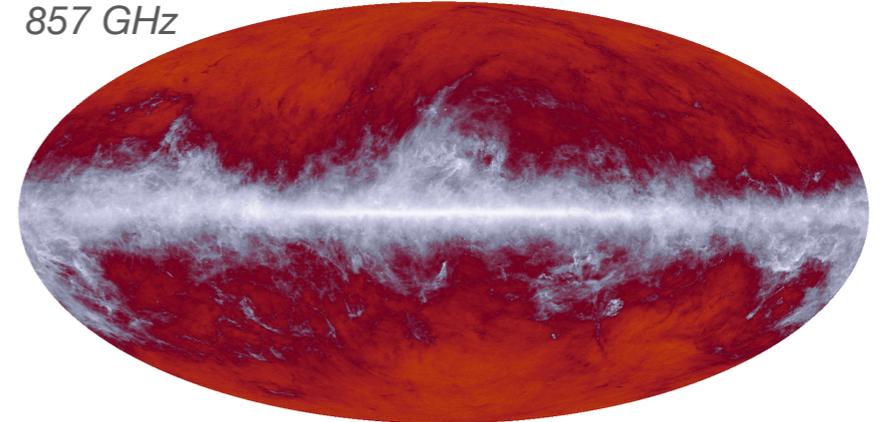
353 GHz



545 GHz



857 GHz

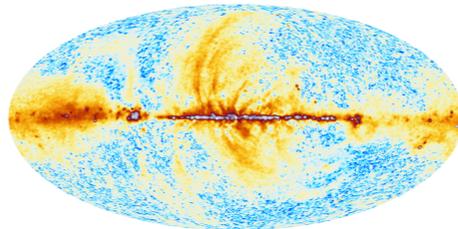
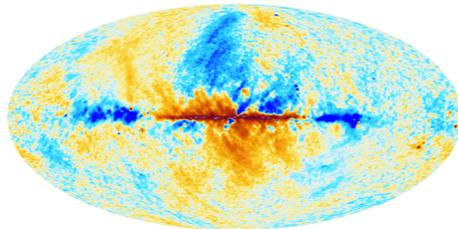
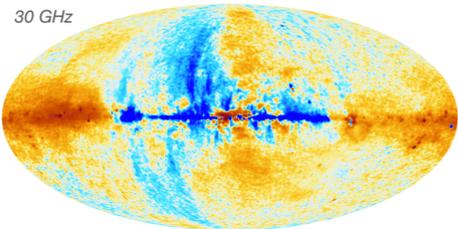


Q

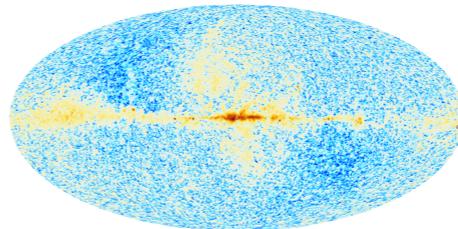
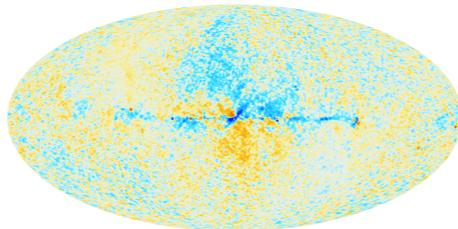
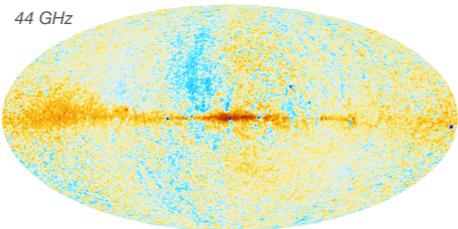
U

P

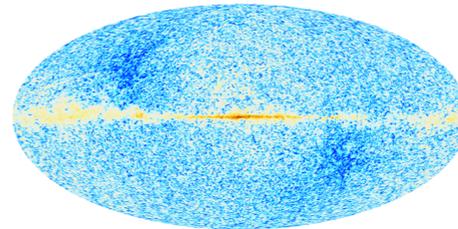
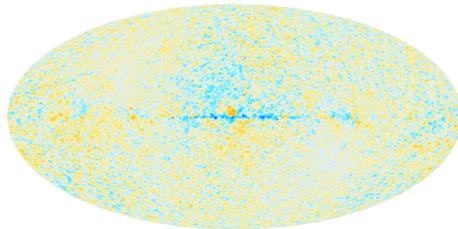
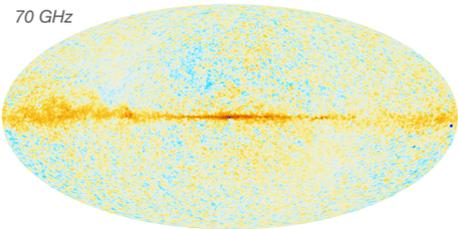
30GHz



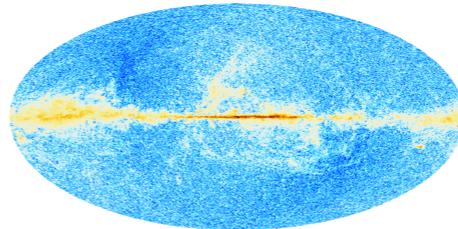
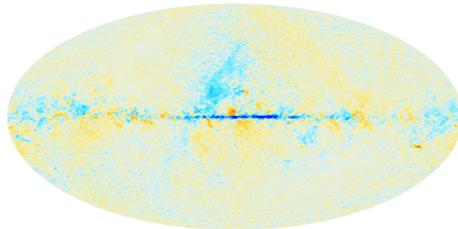
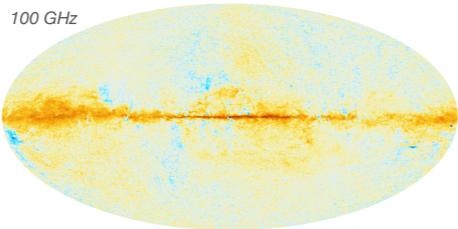
44GHz



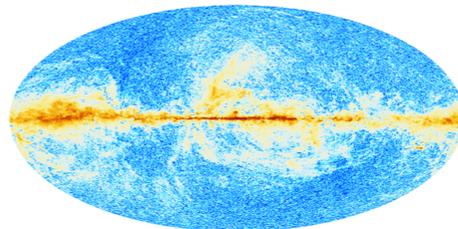
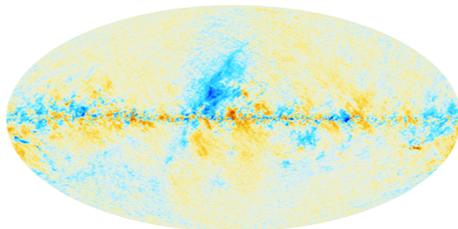
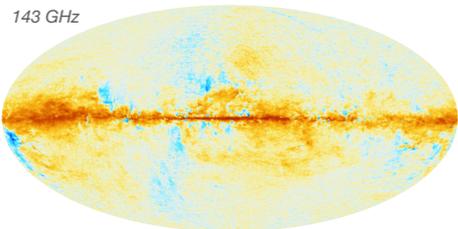
70GHz



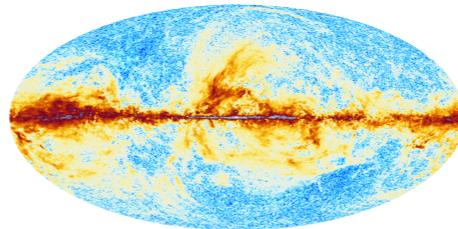
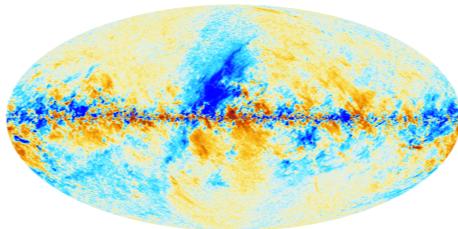
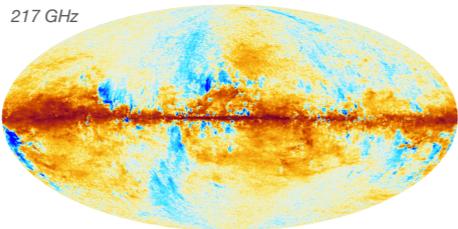
100GHz



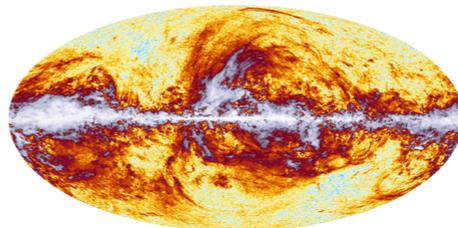
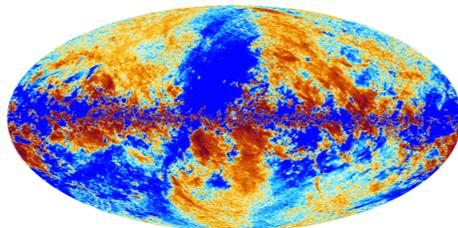
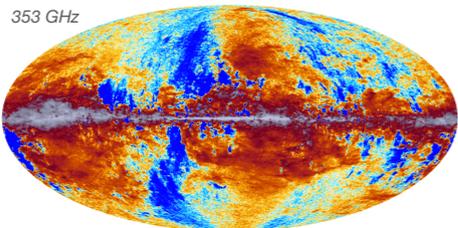
143GHz



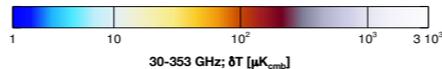
217GHz



353GHz

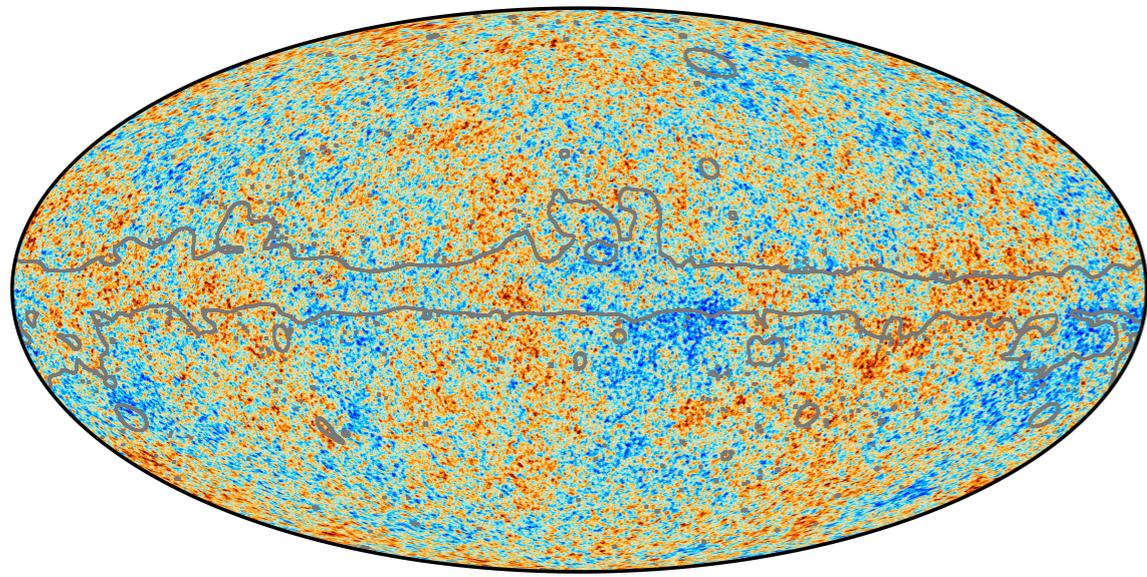


**Plus all-sky
polarization
maps at 6
frequencies**



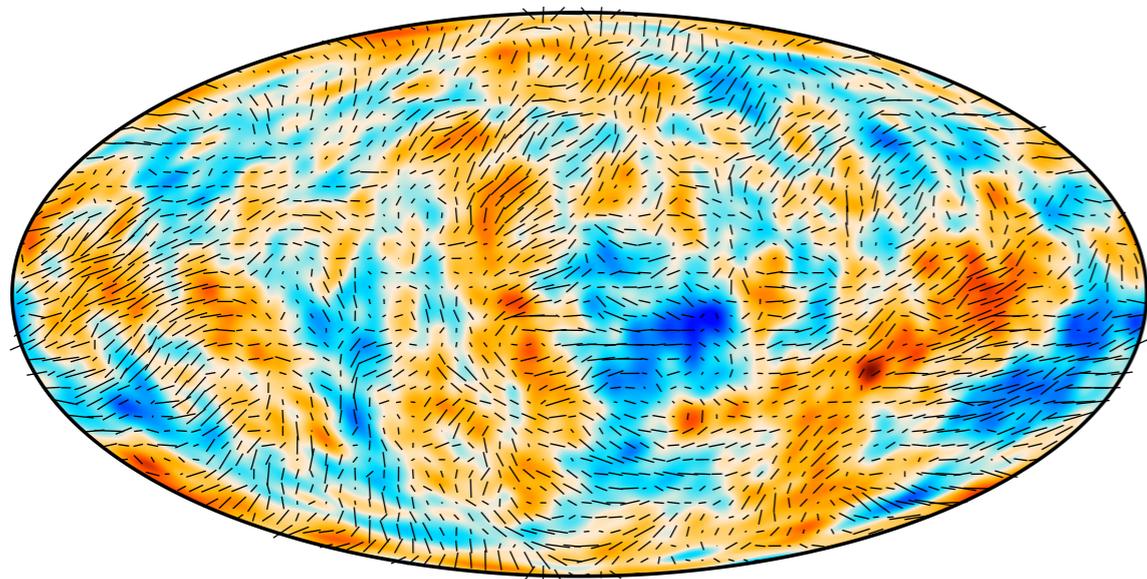
Planck resolution 5'–30'

- **But the whole sky (wide field!)**
- **And large scales well measured (in fact so well measured that we're done, at least in temp.)**



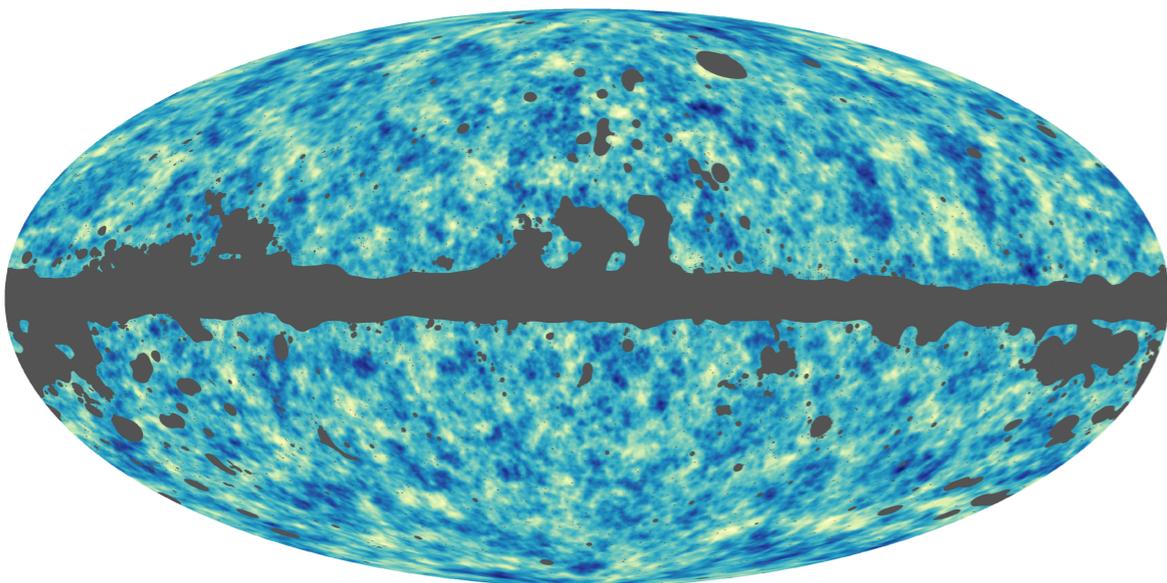
-300 300 μK

CMB temperature



0.41 μK -160 160 μK

**CMB temperature
+ polarization**



-0.0016 0.0016

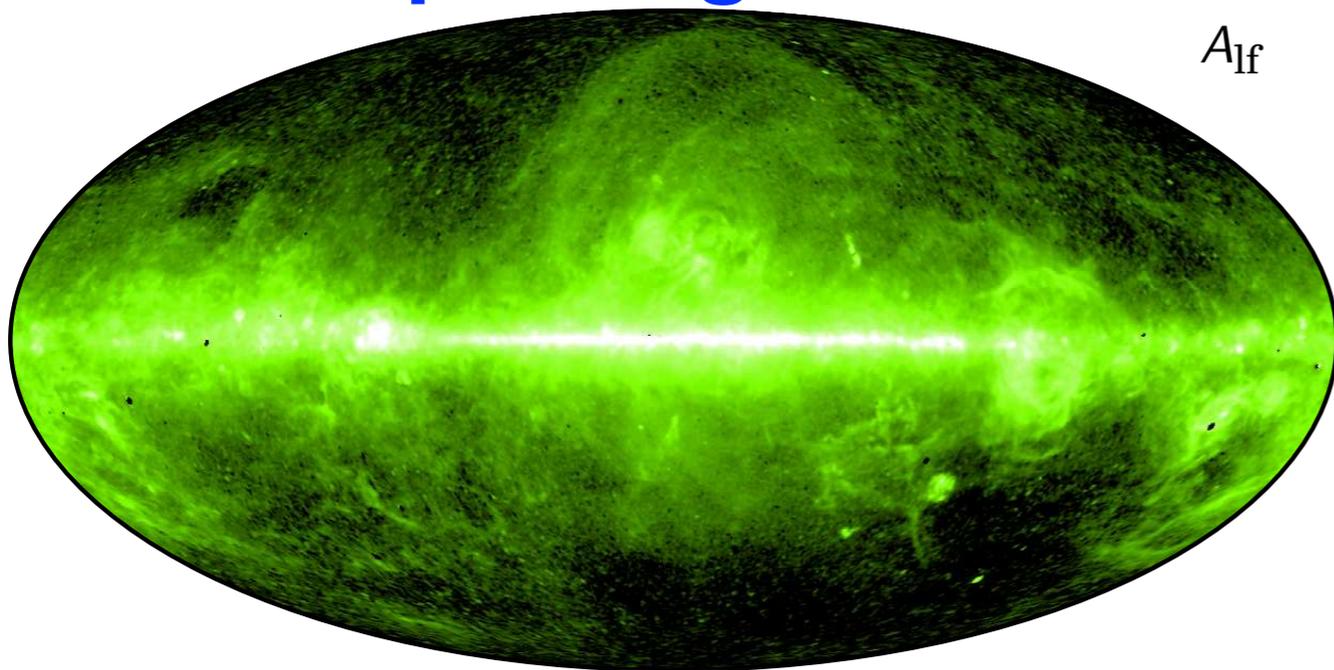
Lensing deflection

But can also study...



...Milky Way from the inside!

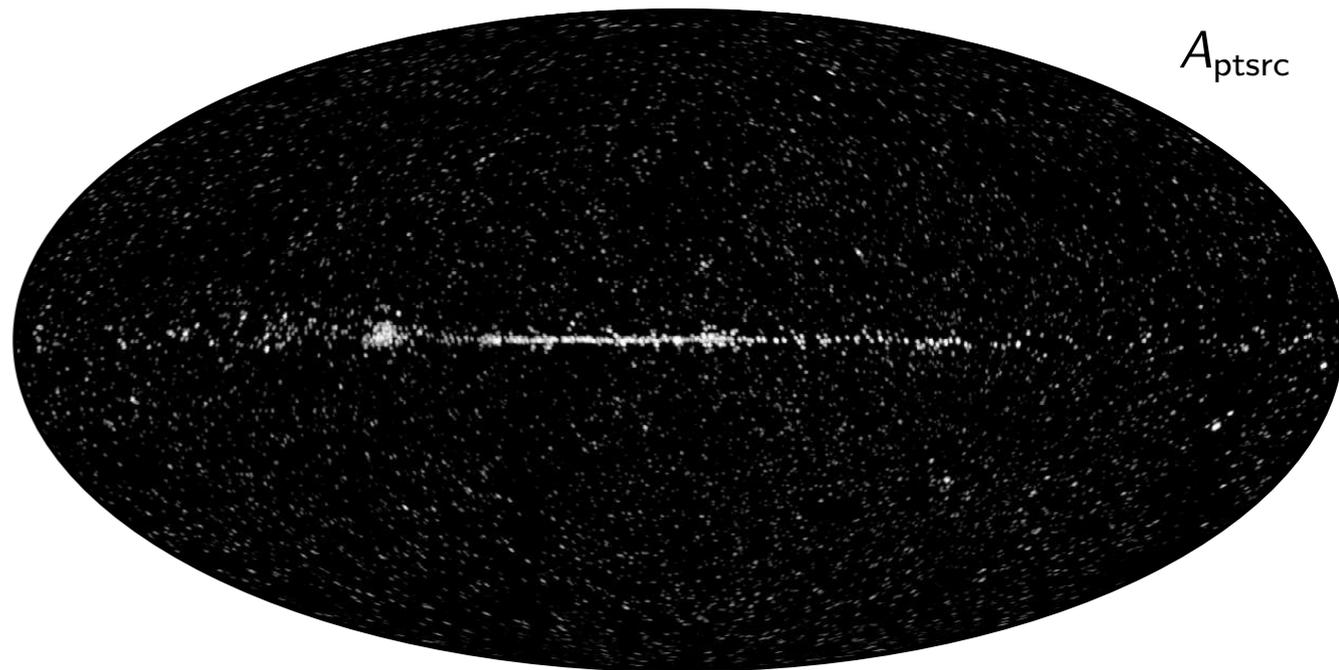
Synch. + free-free + spinning dust



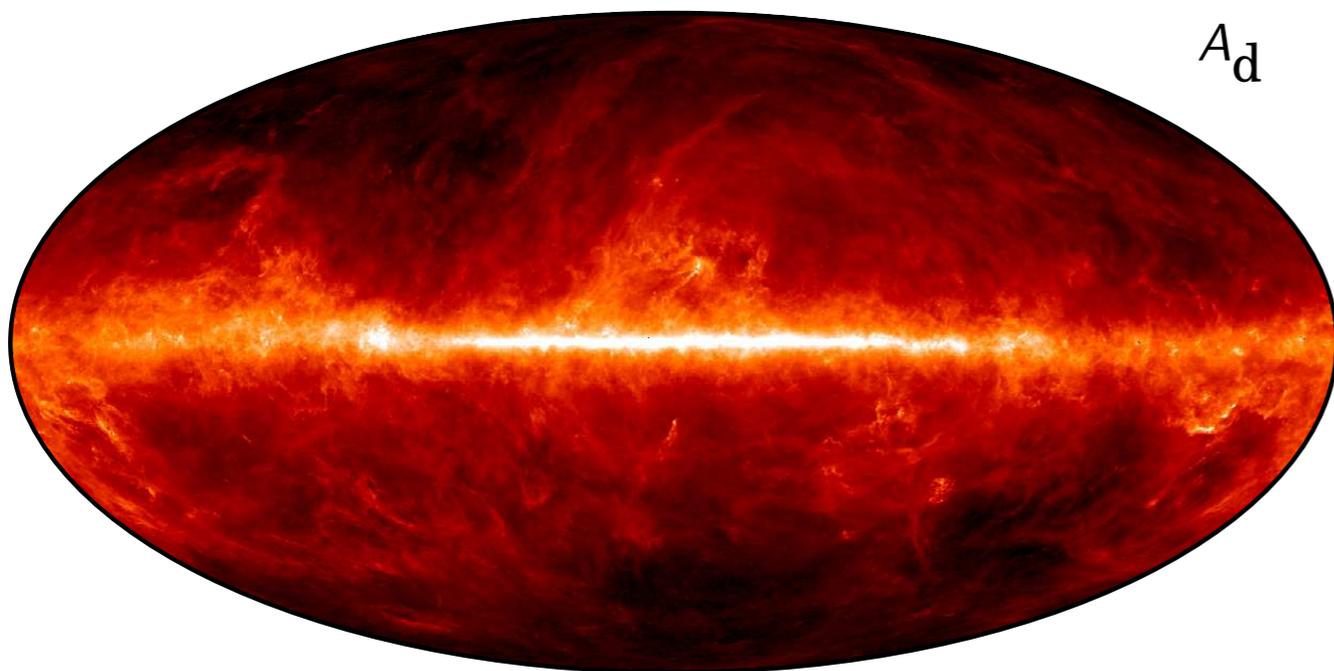
A_{lf}



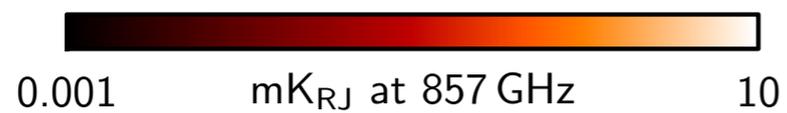
Radio point sources



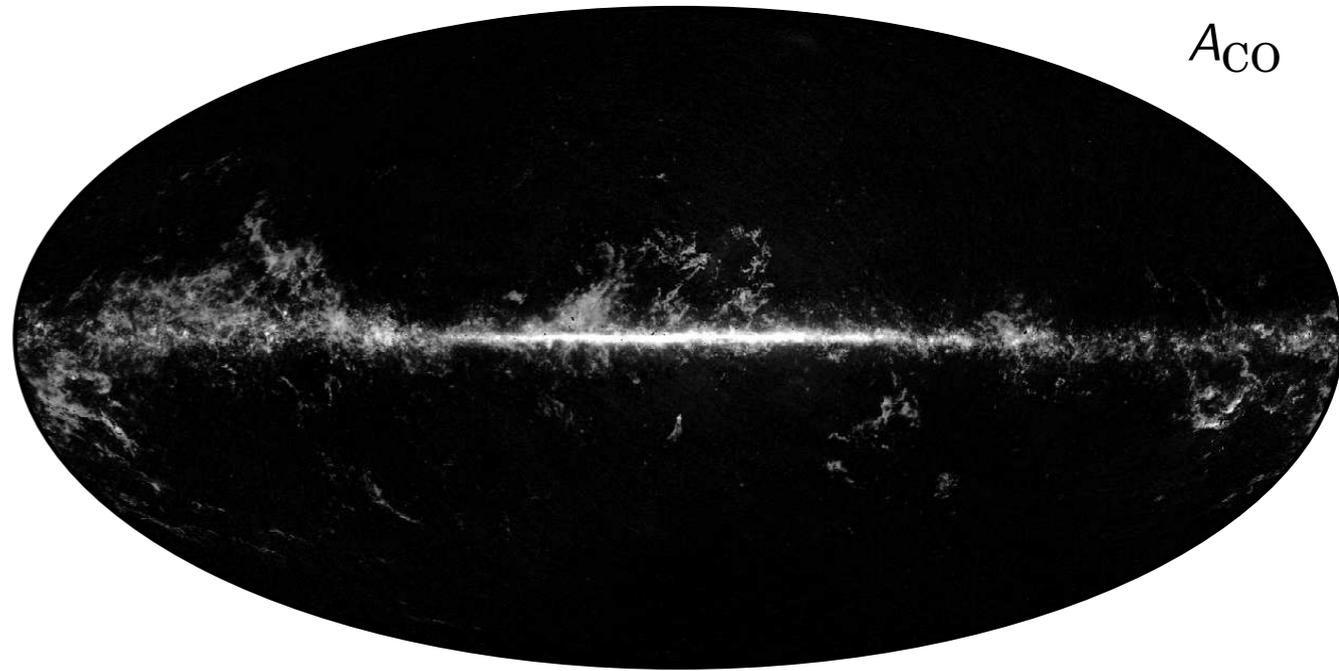
A_{ptsrc}



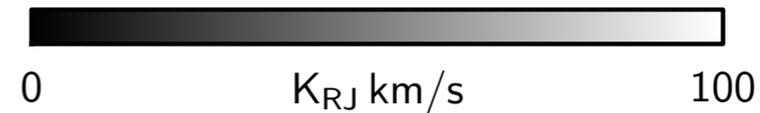
A_d



Thermal dust

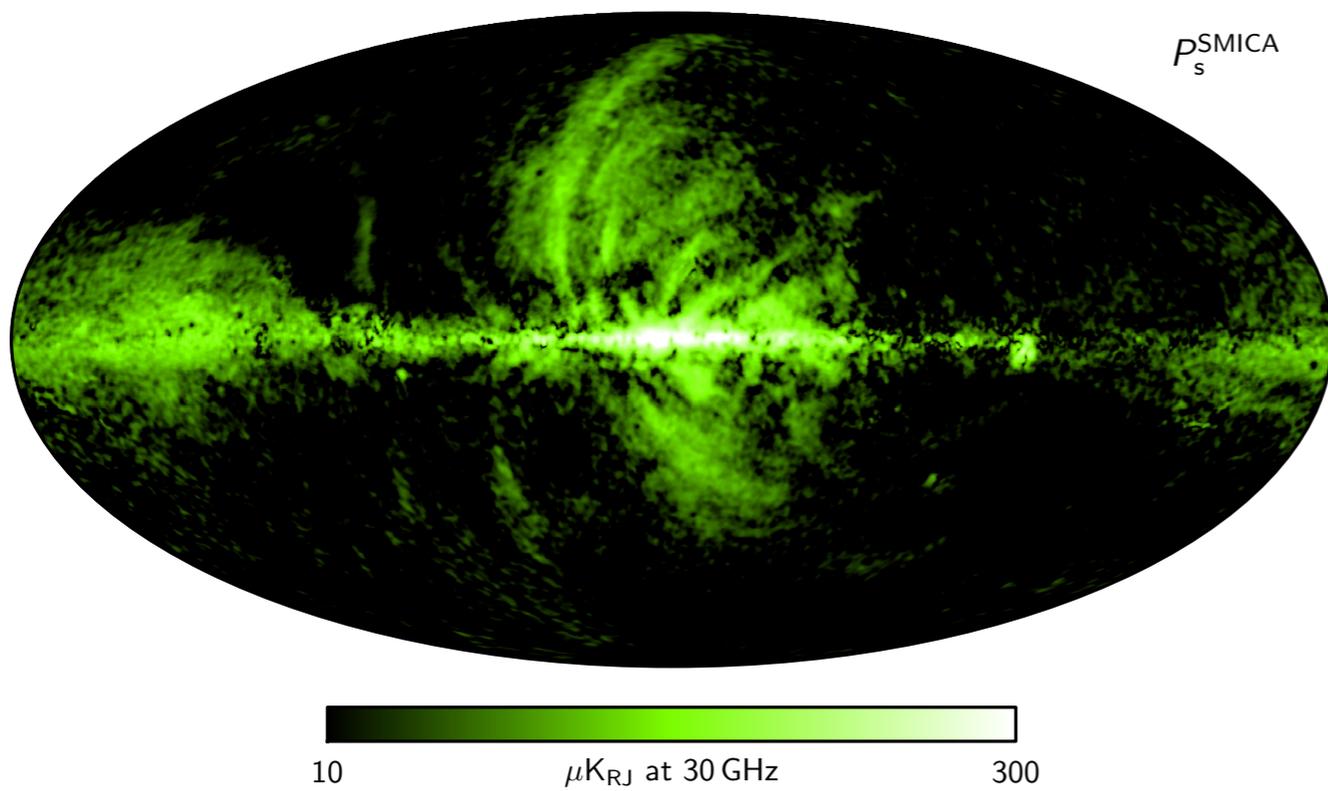
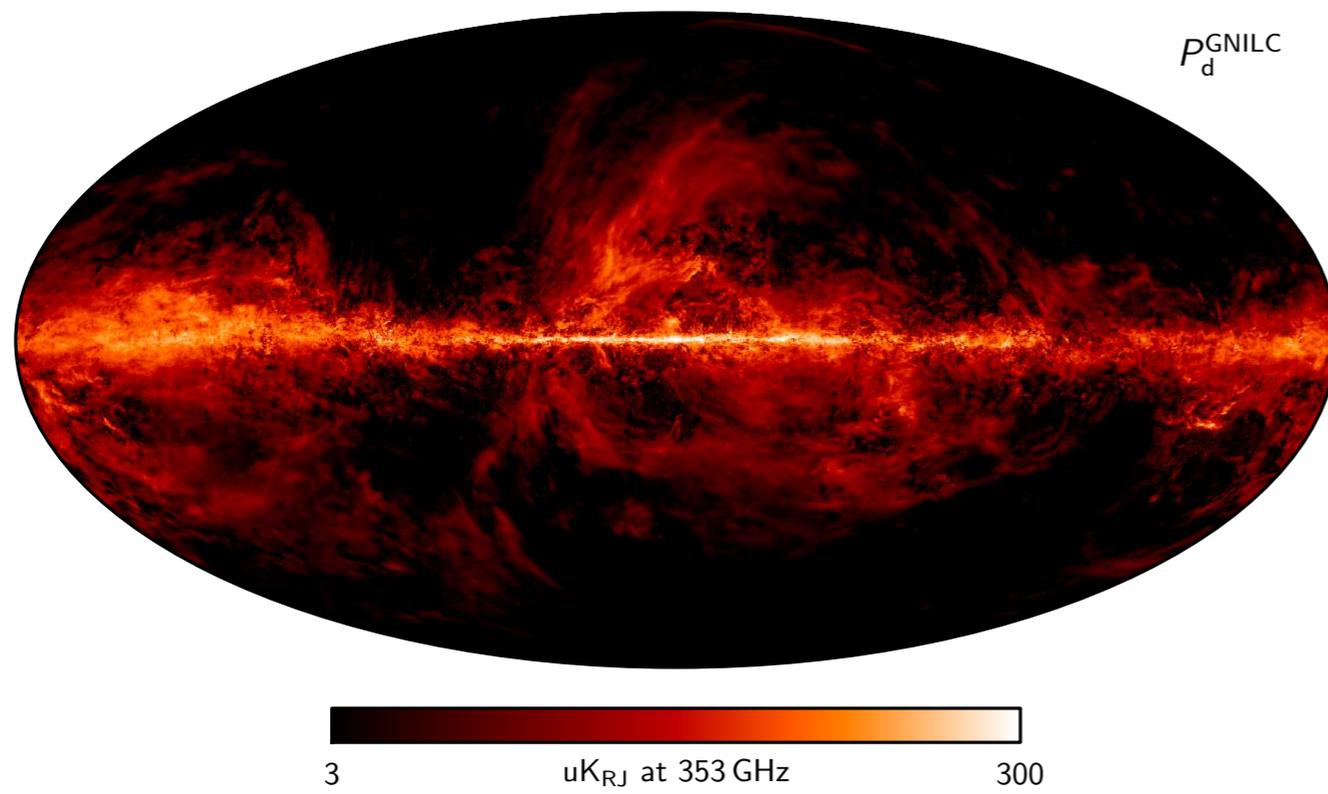


A_{CO}



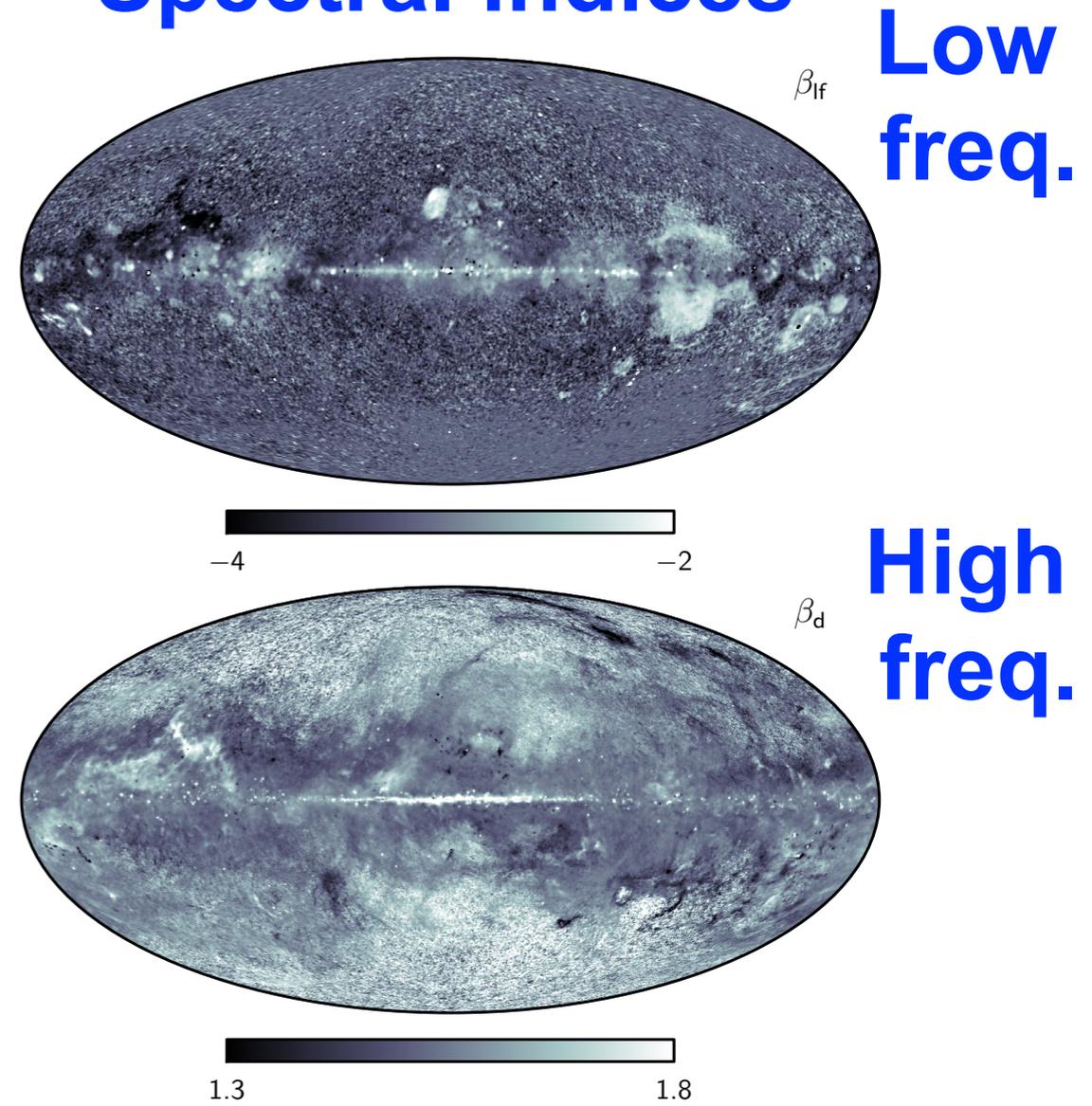
Carbon monoxide

Polarized dust amplitude

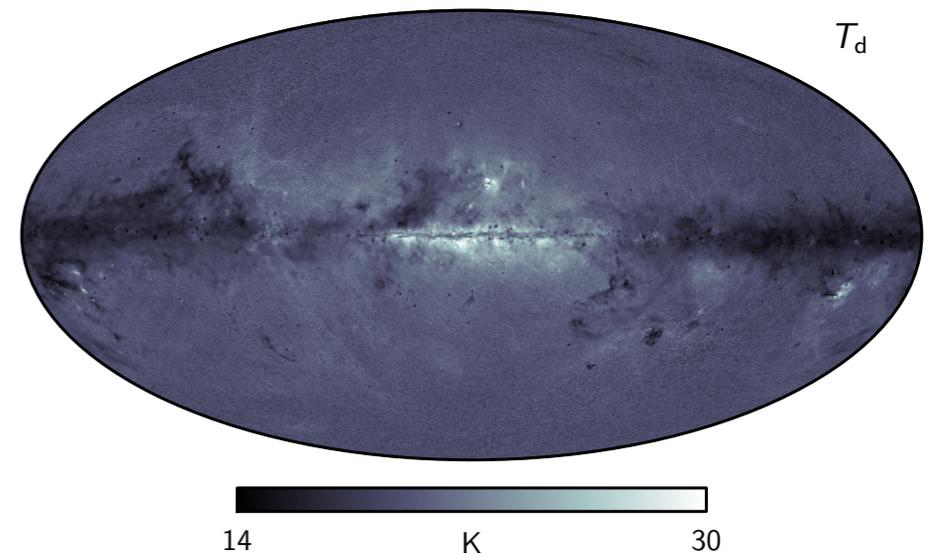


Polarized synchrotron amplitude

Spectral indices



Dust temperature



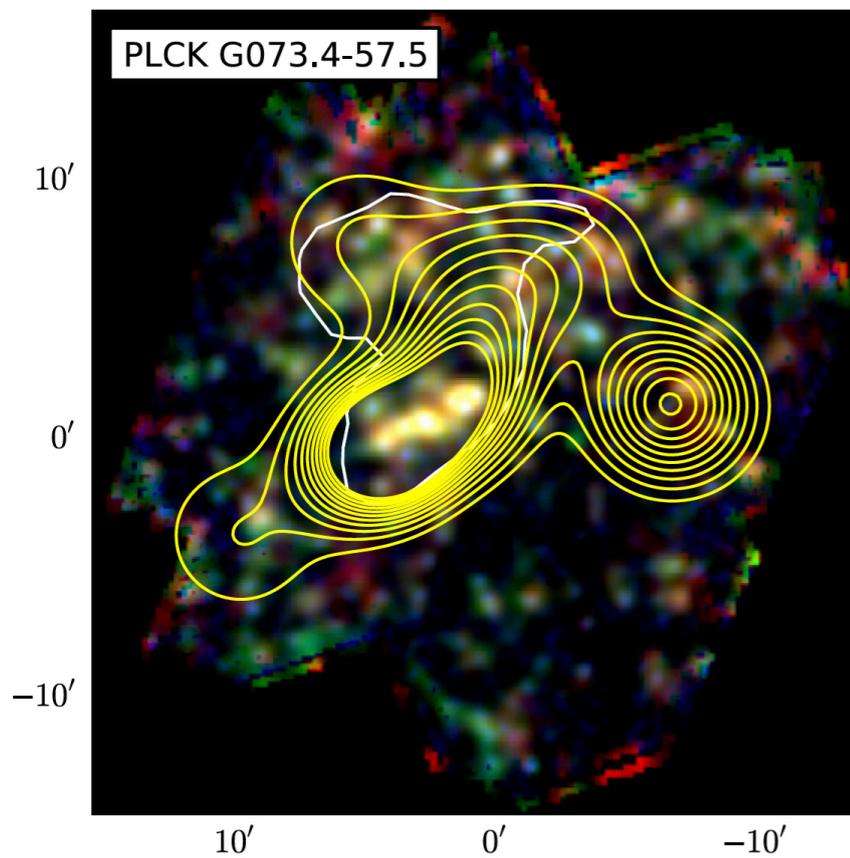
**Plus other extragalactic signals
(of decreasing S/N):**

- **Cosmic infrared background**
- **Sunyaev-Zeldovich y**
- **Integrated Sachs-Wolfe effect**
- **Kinetic SZ effect**
- **...**

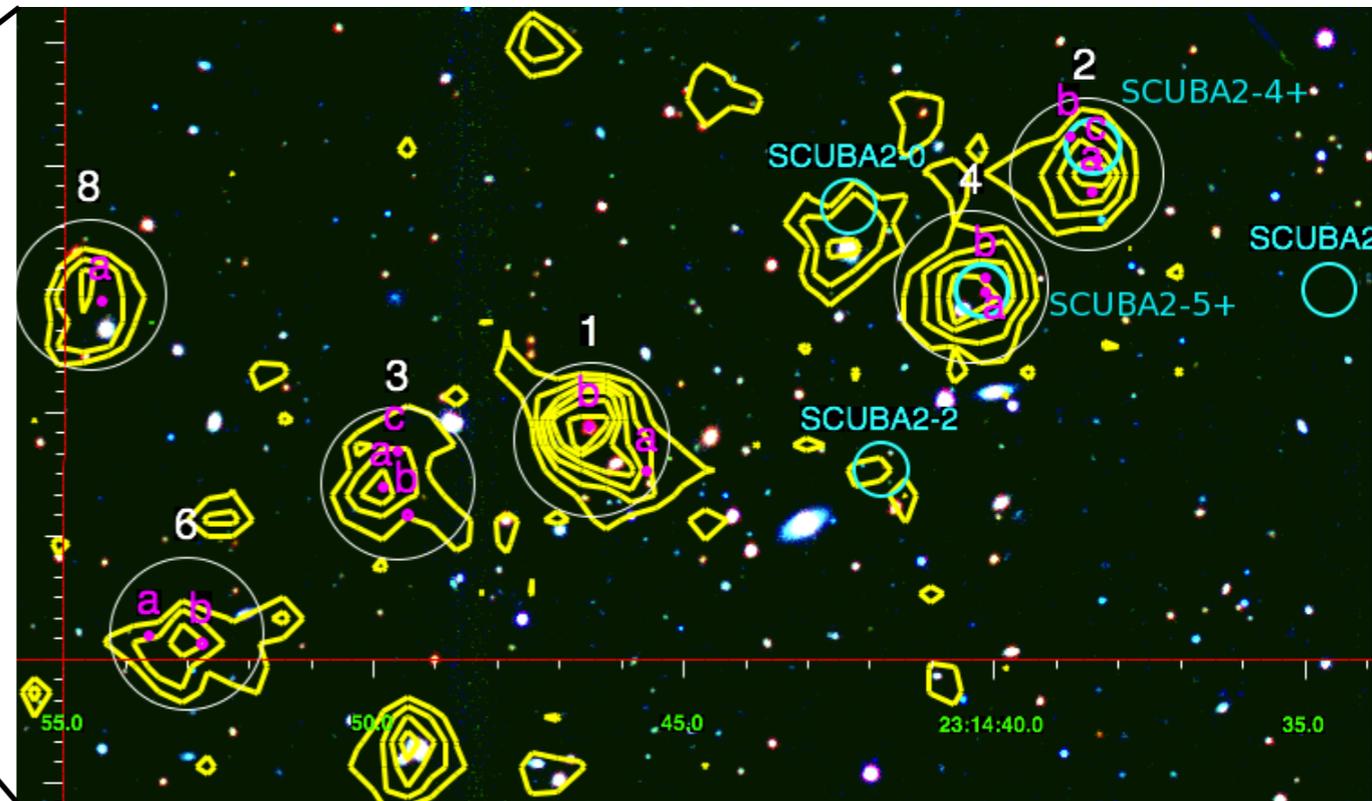
Plus compact sources:

- **Proto-cluster candidates**
- **Strong lenses**
- **...**

Protoclusters



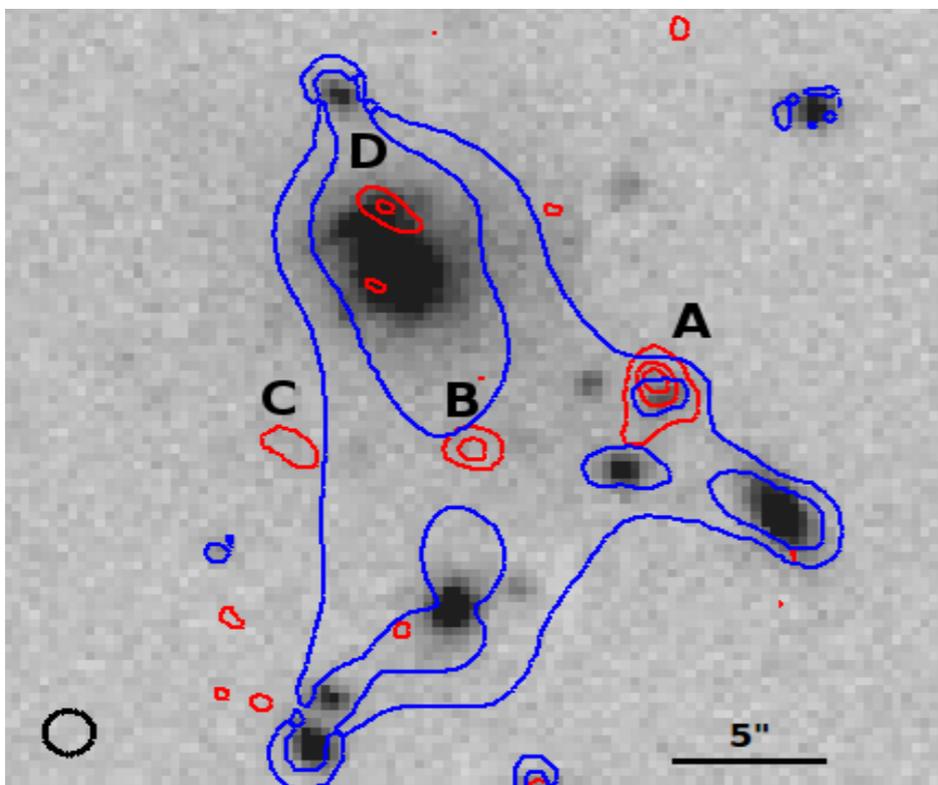
Herschel-SPIRE source overdensity in Planck blob



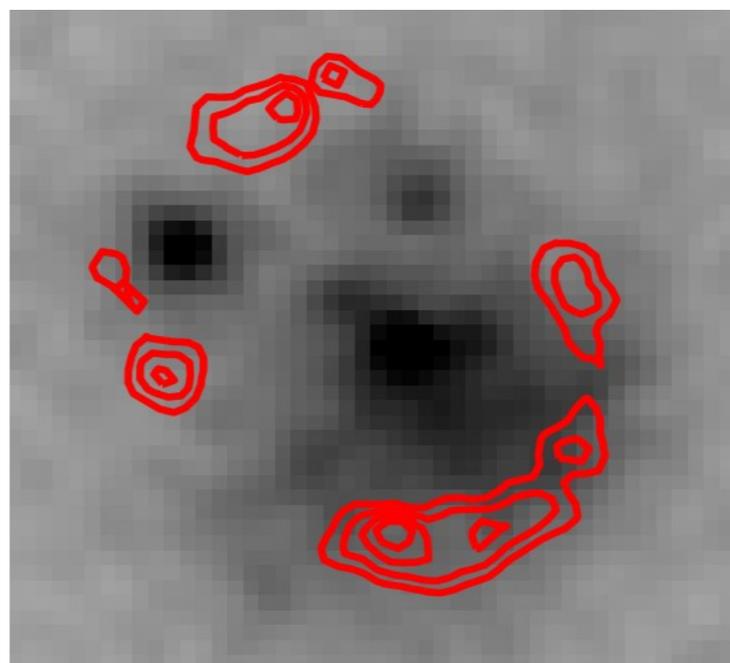
Riley Hill et al.

IRAC+VLT 3-colour image, SCUBA-2 and ALMA follow-up

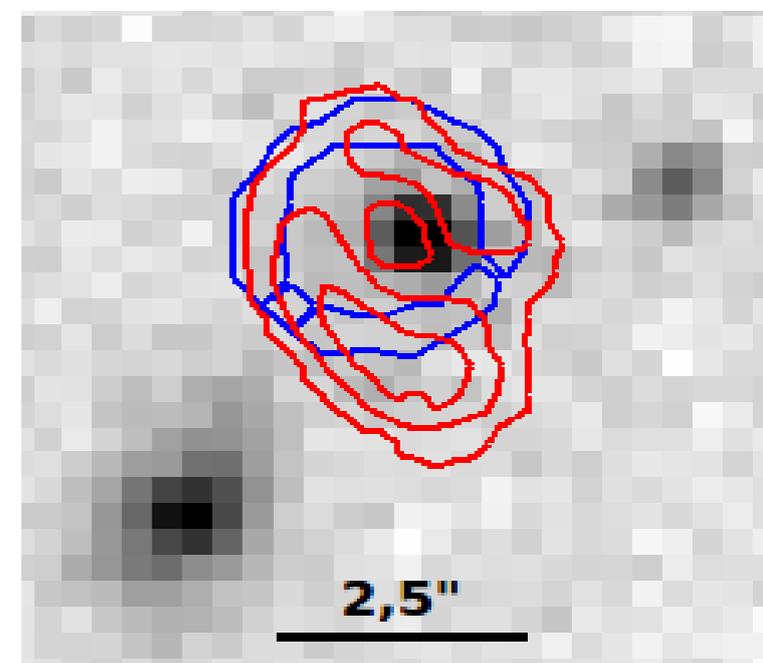
Lenses



Optical + submm contours



Planck-derived sample gives brightest known submm lenses



Ongoing and near future experiments (better resolution/sensitivity/polarization):

- **ACT/ACTPol/AdvACT**
- **SPT/SPTPol/SPT-3G**
- **BICEP/Keck**
- **POLARBEAR/Simons Array**
- **Simons Observatory**
- **...**

Future satellites:

- **Litebird**
- **...**

Litebird: JAXA (Japan) + partners

LiteBIRD is a millimetre-wavelength space satellite telescope being designed and built to map the CMB polarization on degree and larger angular scales. LiteBIRD will be instrumented with 2,622 detectors, spanning observation frequency bands from 40 to 400 GHz, with noise levels about 20 times lower than for the Planck satellite. The mission is led by JAXA (P.I.: Masashi Hazumi) and is presently in Phase A1. The Canadian-designed DfMUX bolometer readout system is the baseline for the mission.

LiteBIRD

LiteBIRD - Lite (Light) Satellite for the Studies of B-mode Polarization and Inflation from Cosmic Background Radiation Detection

「学術会議マスタープラン2014」の重点大型研究計画の一つ

ミッション 代表的インフレーション宇宙モデルを全て検証する

- 戦略**
- 緊急度が高く宇宙でのみ可能な観測に的を絞り、大きな成果を世界に先駆け獲得
 - 明確な目標(原始重力波強度パラメータ r を誤差0.001以下で測定)に依る設計
 - 小型光学系と多色焦点面により重量・コスト減
 - 地上観測プロジェクトによりサイエンスの成果を出しつつ技術実証

衛星システム

光学系

- 回転半波長板による変調
- クロスドラゴン方式の主鏡、副鏡(直径約60cm)を4ケルビンに保持



- バスシステム**
- データ取得系
 - 制御系
 - 通信系

- スターリング冷凍機
- ジュールトムソン冷凍機
- 断熱消磁冷凍機(X線天文学のノウハウ活用)

観測

- 全天CMB偏光サーベイ
- スピンレート: 0.1rpm
- 50~320GHzの間で6バンド観測
- 角度分解能: 30分角@150GHz

主な仕様

- 観測期間: 3年(目標値)
- 軌道: L2 (ただしLEOと比較検討中)
- 重量: 約1t
- 電力: 約2kW

- **Launch ~2026**
- **Study for Canadian involvement underway**
- **Ask Matt, Renee, Douglas or Dick**

CMB-S(tage)4

Primordial B modes: Simon Foreman already discussed this

- **Constrain r at the 0.001 level**
- **Constrain (neutrino) N_{eff} at 0.05 level**
- **Survey ~half sky at mm to cm**
- **20-270GHz, multiple channels**
- **~500,000 detectors**
- **2 distinct sets of telescopes**
~0.5m and ~6m (Chile and SP)
- **~5% sky survey really deep**
- **+ ~40% sky survey just deep**

Commissioning 2025?

This will happen!

Why should I care about CMB-S4? (if I think r and N_{eff} are boring!)

- 40% of the (southern) sky
- Stokes I , Q , U
- SZ, lensing, ...
- Galactic polarization + sources
- Cadence ~every couple of days
- 1.5' beam at 150GHz (scales like $1/\nu$)

- Variable and moving sources?

Report writing for (US) decadal survey

- **“Legacy archive” working group coordinators, DS + Lindsey Bleem**
- **Need help thinking about the non-cosmology stuff!**

CMB Lensing Map

- Legacy product will be a lensing map(s) for 40% of the sky covered by the wide area survey
- Excellent overlap with e.g., LSST, DESI, eRosita, etc
- Improved Cosmological Constraints via cross-correlations with other LSS Surveys (e.g., Doux+18, and others)
- Calibration of Systematic Errors in other surveys (e.g., Schaan+17, and others)
- High-redshift mass calibration via CMB halo lensing (clusters, AGN, other sources; e.g., Raghunathan+17, Melin+15, and others)

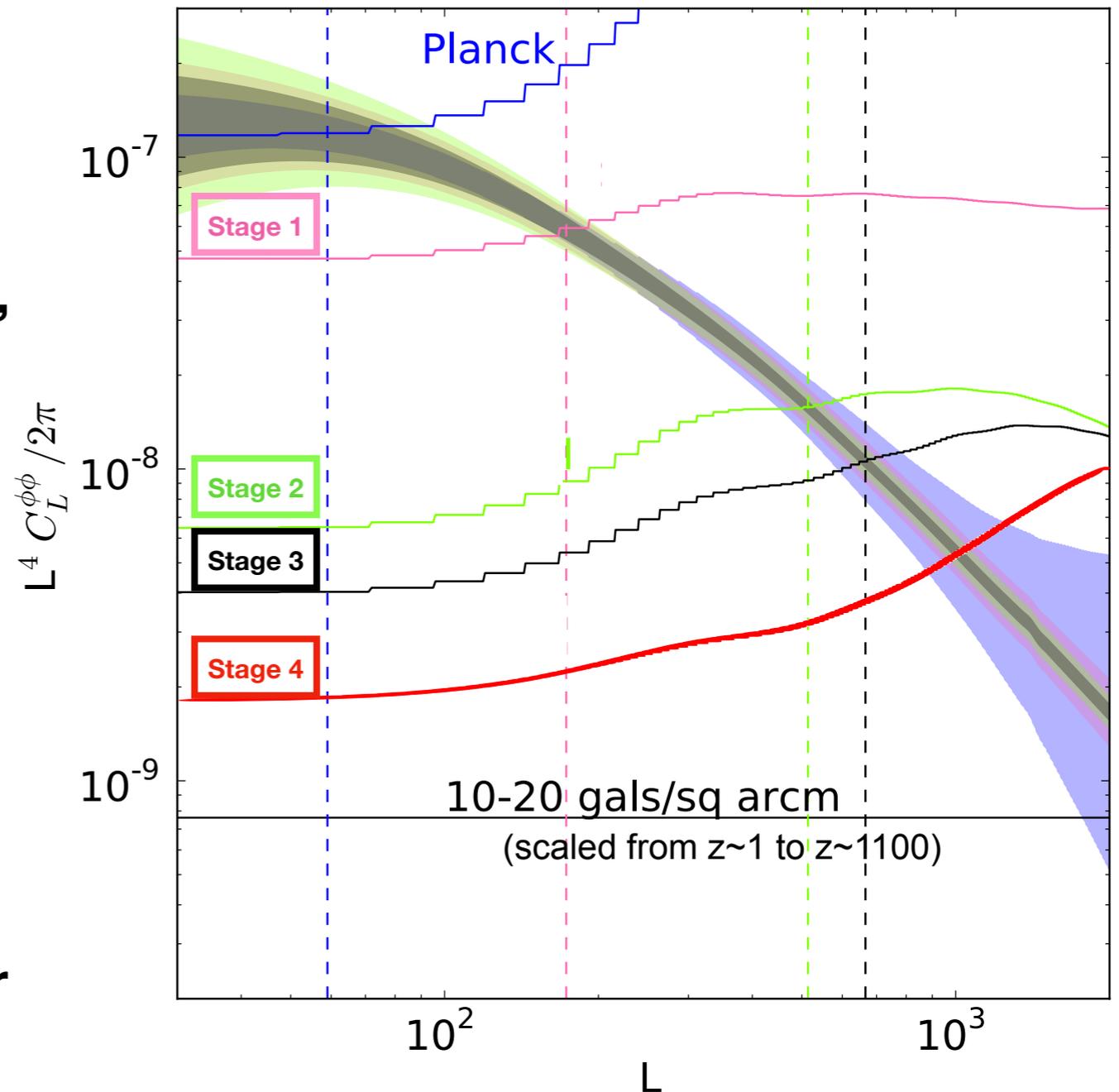
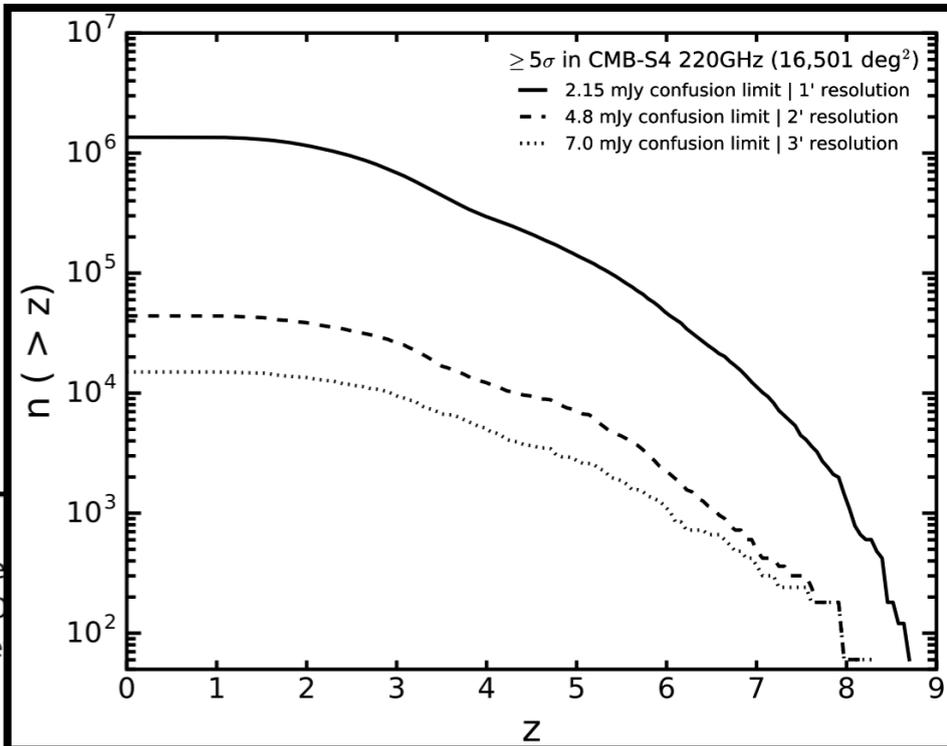
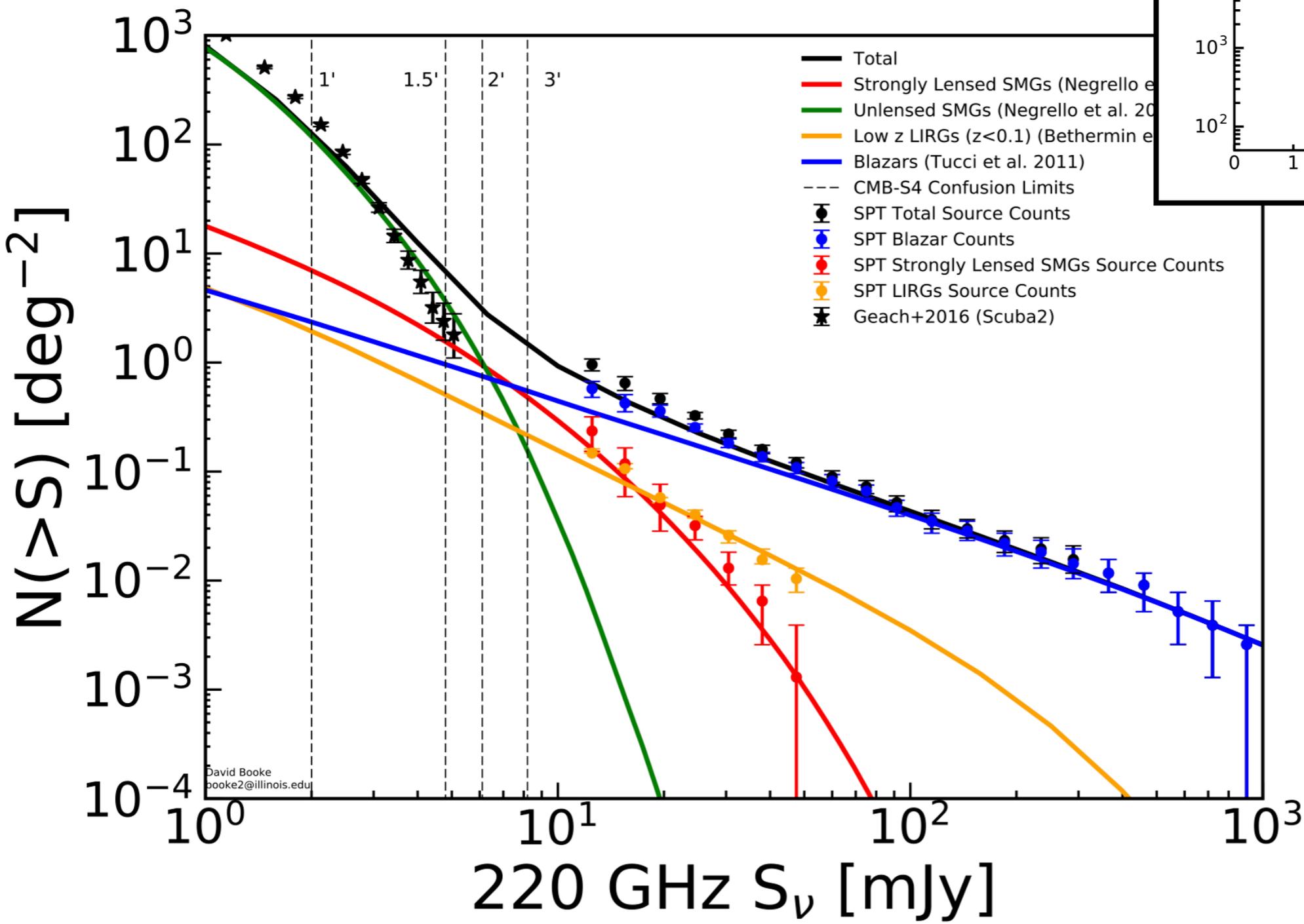


Figure G. Holder; S4 noise curve J. Meyers

Extragalactic Sources

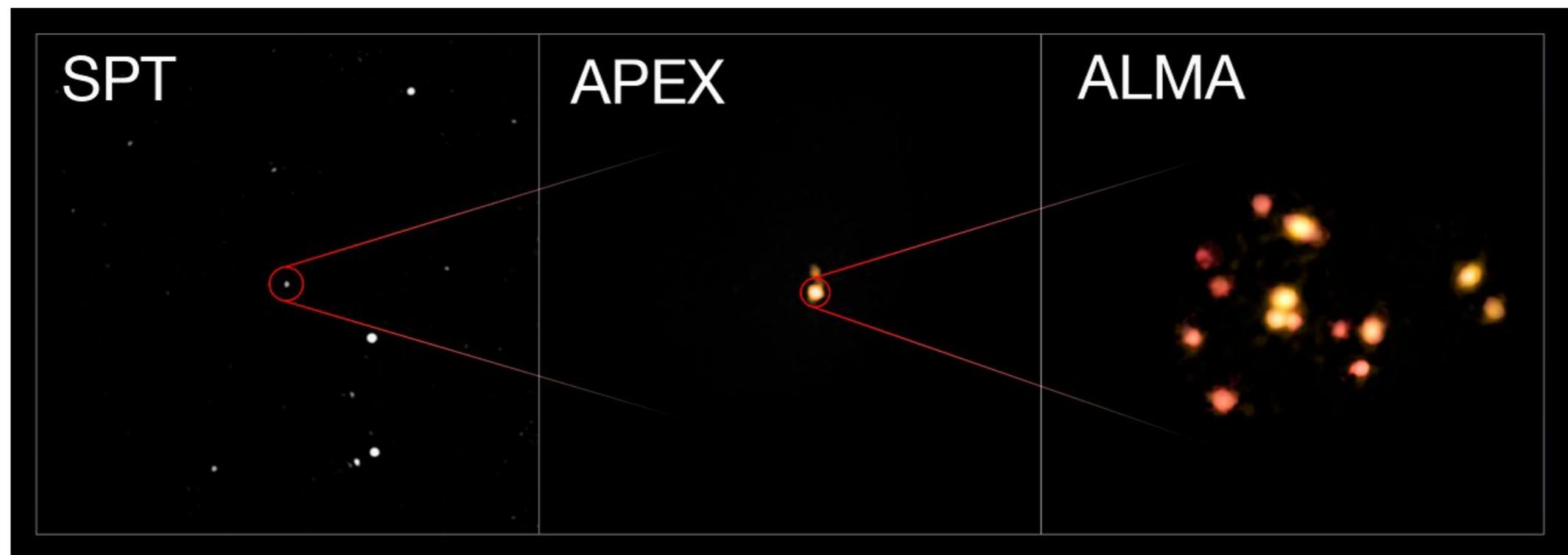
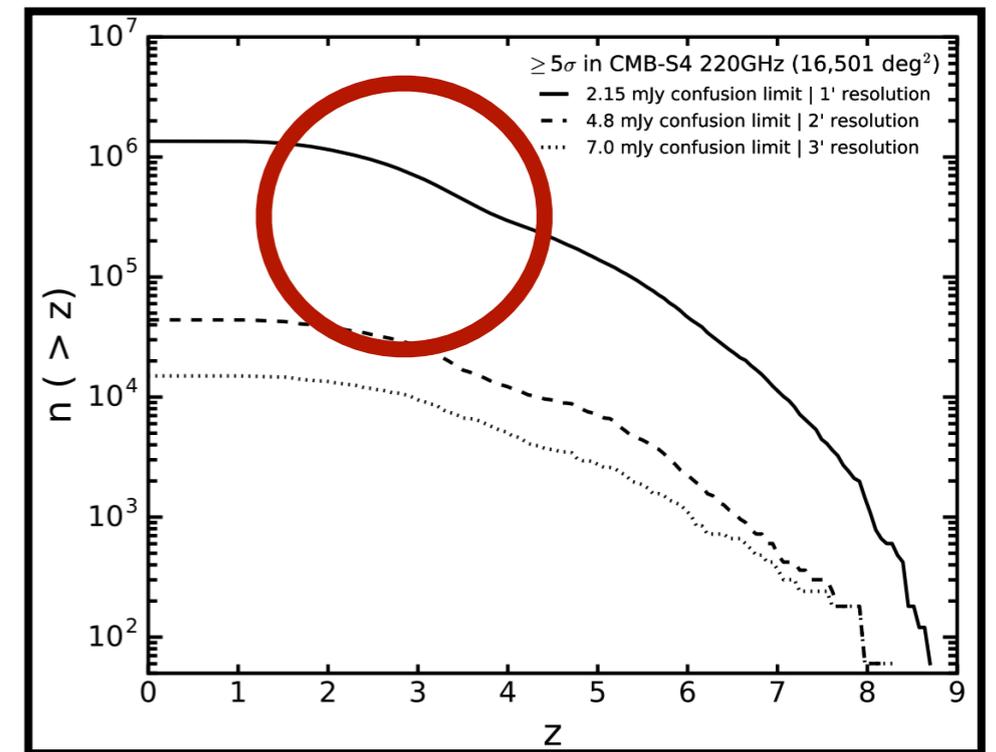


David Booke
booke2@illinois.edu

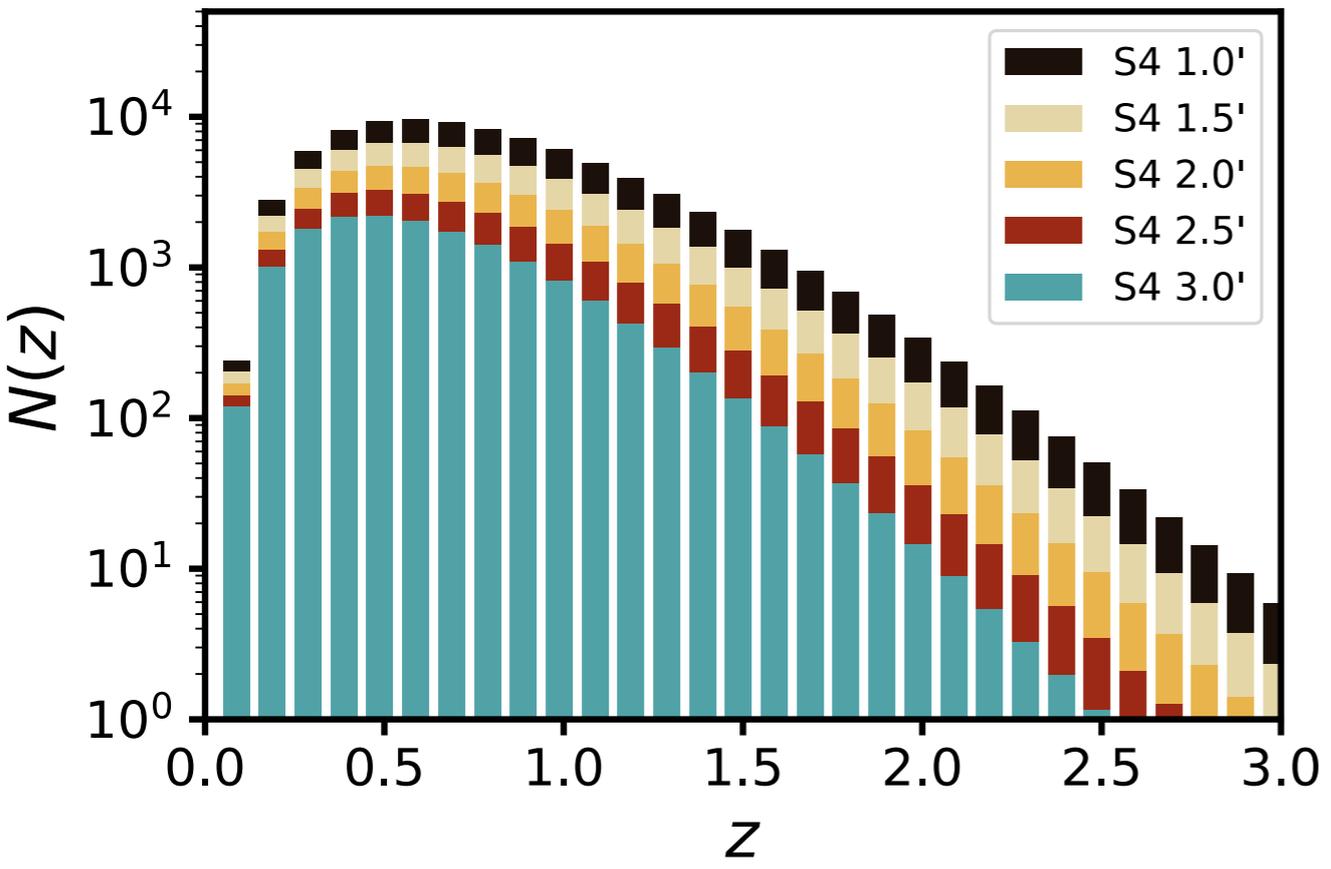
Figures from Joaquin Vieira

Extragalactic Sources

- Some of the DSFGs identified by CMB-S4 will be protoclusters (high redshift progenitors of galaxy clusters)
- Planck's "PHz" catalogue contains about 2000 "red peaks" in the CIB, over about a quarter of the sky
 - >200 followed up with Herschel, but we still don't really know what these are
 - there are hence no good theoretical predictions for their abundance
- CMB-S4 will detect many more

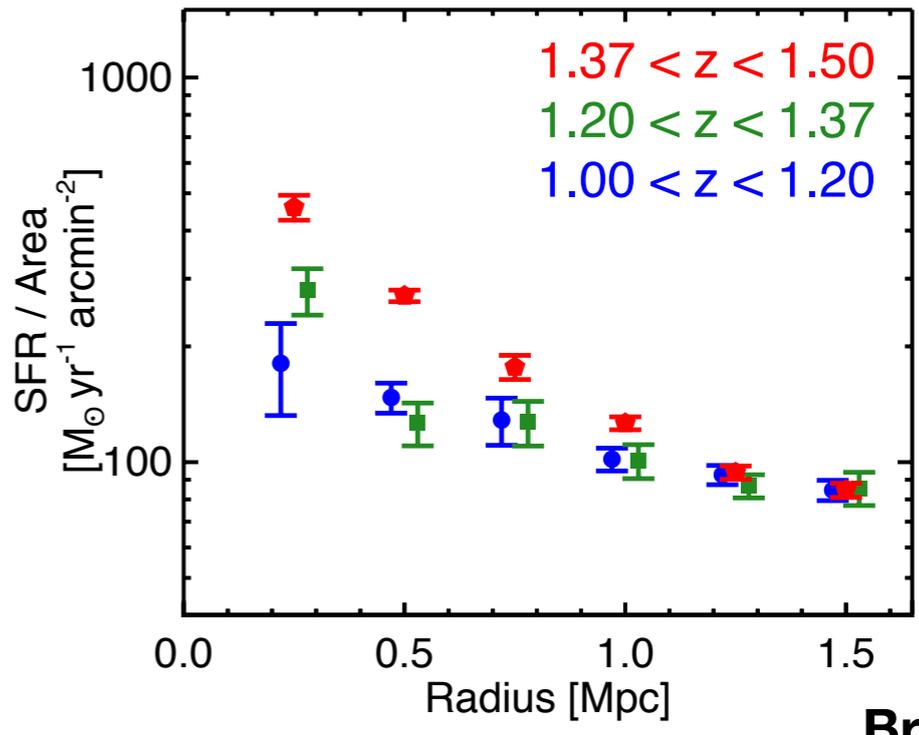


Galaxy Clusters: Cosmological and Astrophysical Implications

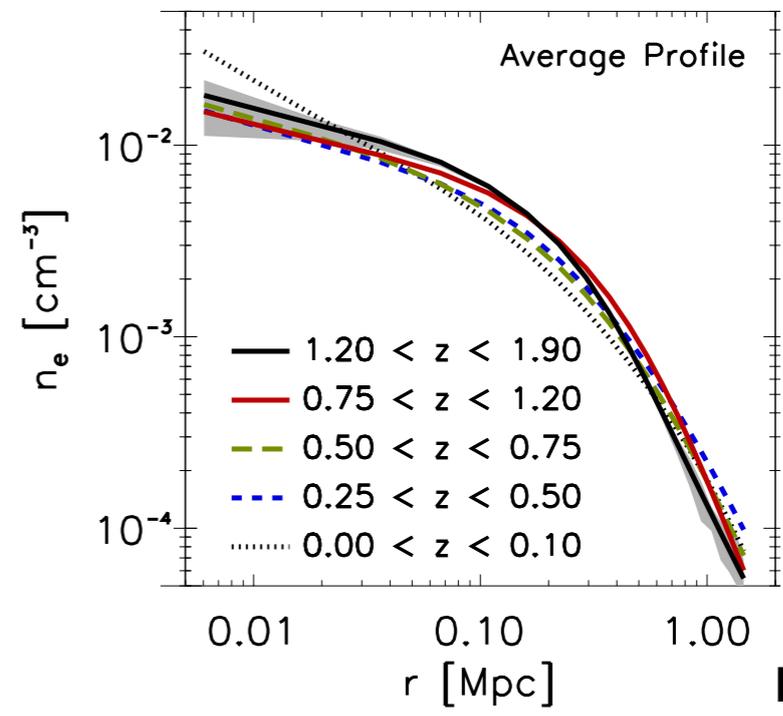


Madhavacheril, Battaglia, Miyatake 2017

“The number of massive galaxy clusters could emerge as the most powerful cosmological probe [if systematics can be controlled]” - DOE Cosmic Visions Dodelson+1604.07626



Brodwin+13

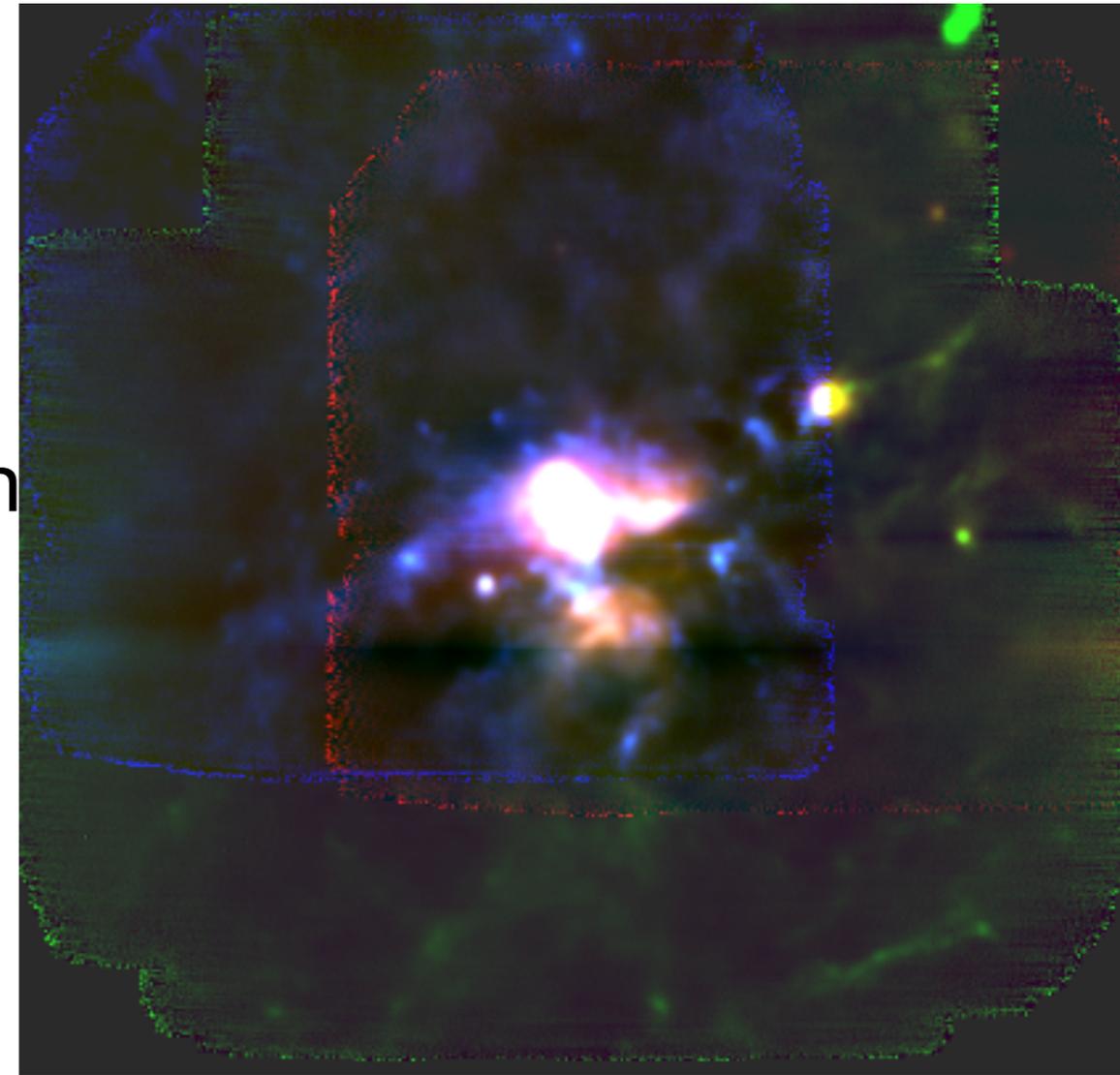


McDonald+17

Wonderful Synergy with LSST, Euclid, WFIRST, eRosita, Athena, Lynx and more!

Galactic Sources

- Planck (and earlier ground-based surveys) catalogue ~150 CO clouds at $|b| > 30^\circ$
- Planck also found hundreds of “cold cores” at high latitudes (some in known GMCs, others unknown)
- H-ATLAS discovered some sources associated with stars, including a few debris disk candidates
- As well as extracting Galactic sources in CMB-S4 maps, we can cross-correlate with stellar (and other Galactic) catalogues
- Not the highest priority - but does allow us to connect with another community of astronomers



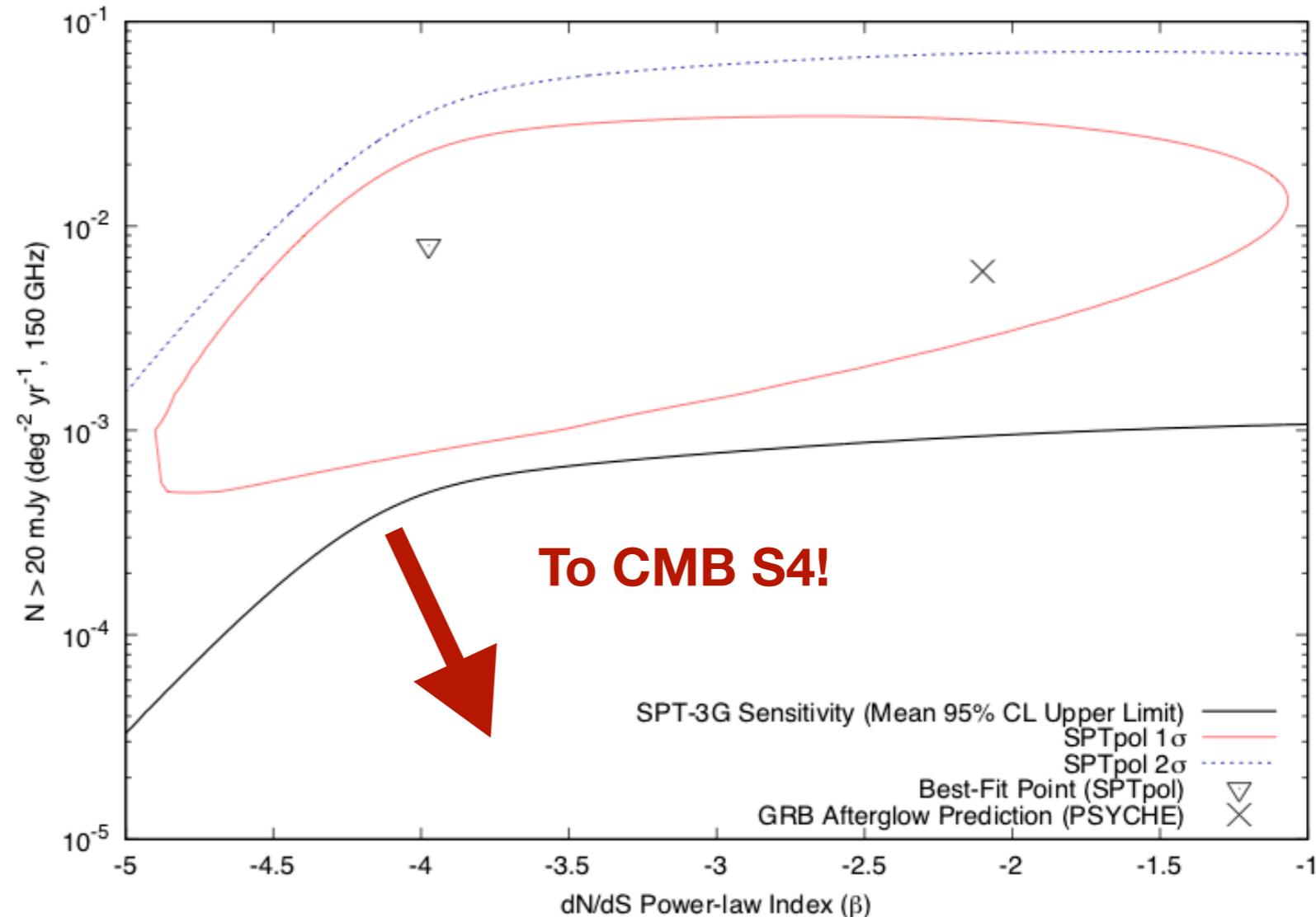
RCW 38 HII region
(90, 150, 220 GHz; rgb)
Credit: K Schaffer, SPT-SZ

HELP WANTED!

Transients - GRBs and FRBs

- Orphan GRB afterglows (generic prediction of GRB models, but none yet detected)
- New probe of the Epoch of Reionization! CMB S4 should be able place interesting constraints on some models of GRBs from population 3 stars at $z \sim 20$
- FRBs challenges include detector glitches, Cosmic Rays

Figure by Nathan Whitehorn



Forecasting Machinery in place, need straw cadence to run projections for CMB-S4.

Near Earth Objects

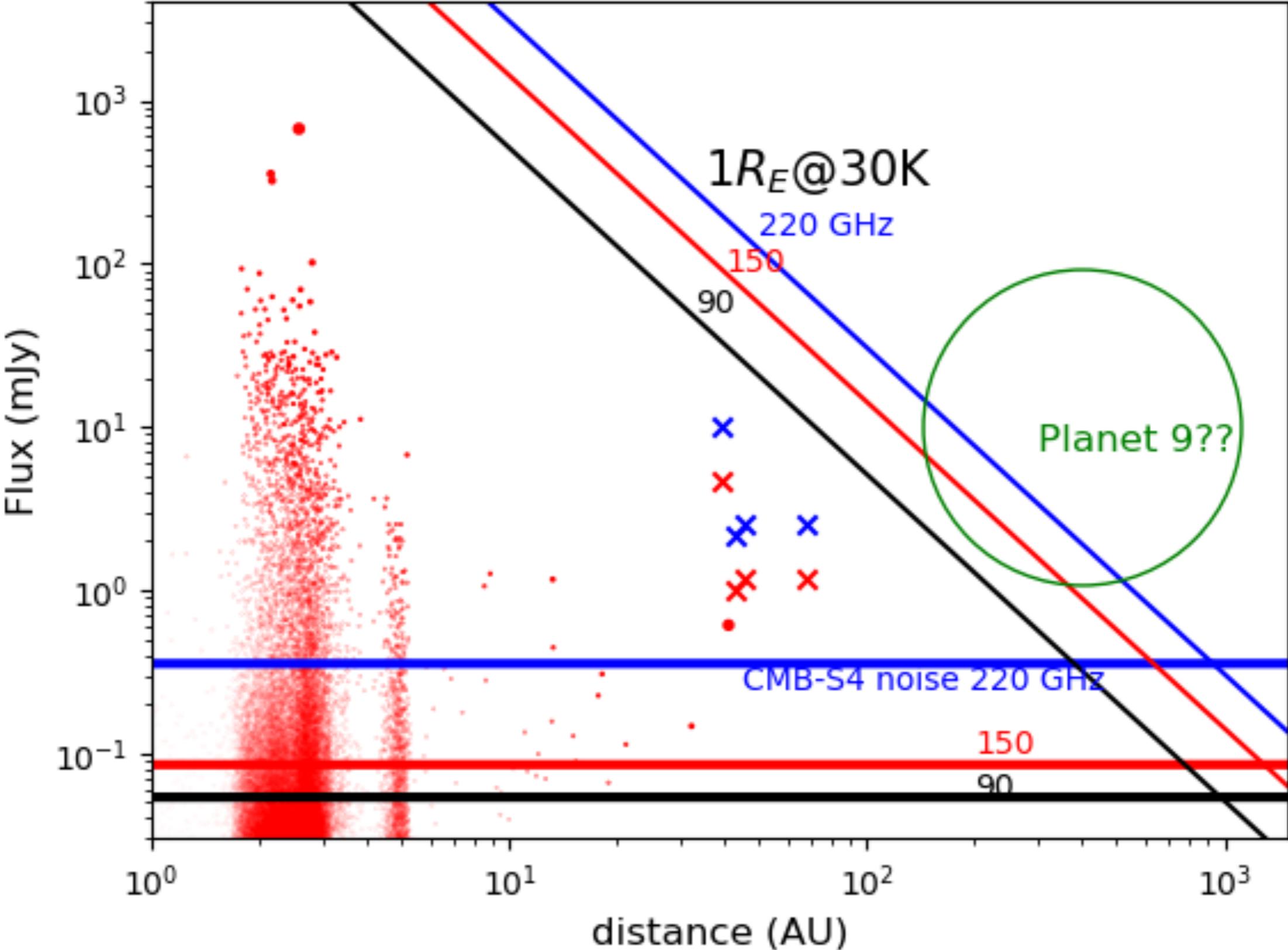


Figure by Gil Holder

Galactic sources

- In addition to Galactic polarization studies, CMB-S4 will detect sources within our Galaxy
- Mostly CMB-S4 people won't bother with Galactic sources - but they'll be seen anyway!
- Planck (and earlier ground-based surveys) catalogued ~150 CO clouds at $|b| > 30^\circ$
- Planck also found hundreds of "cold cores" at high latitudes (some in known GMCs, others unknown)
- H-ATLAS discovered some sources associated with stars, including a few debris disk candidates
- As well as extracting Galactic sources in CMB-S4 maps, we can cross-correlate with stellar (and other Galactic) catalogues

Proto-clusters

- SZ sources are already part of the CMB-S4 science case - hence a full catalogue + diffuse map
- But there are also “proto-clusters” with a different spectral signature - like redshifted clumps of DSFGs
- Planck’s “PHz” catalogue contains about 2000 “red peaks” in the CIB, over about a quarter of the sky
 - >200 followed up with Herschel, but we still don’t really know what these are
 - there are hence no good theoretical predictions for their abundance
- CMB-S4 will presumably detect many more with its smaller beamsize
- Perhaps CMB-S4 can bridge the gap between genuine clusters and proto-clusters?